International School on Applications with the Newest Multi-spectral Environmental Satellites

held in Brienza from 18 to 24 Sep 2011





Dottorato di Ricerca in Metodi e Tecnologie per il Monitoraggio Ambientale



Paul Menzel, Valerio Tramutoli, & Filomena Romana

Su 1 – 4 pm	Lecture 1	Instruments and Orbits
Mo 9 – 12 am	Lecture 2	Radiative Transfer in the Earth Atmosphere
	Homework 1 (due	e Thursday)
Mo 2 – 5 pm	Lab 1	Planck Function and Intro to Hydra
Tu 9 – 12 am	Review	Lecture 1 & 2, Lab 1
	Lecture 3	Spectral signatures from Earth
Tu 2 – 5 pm	Lab 2	Interrogating Multispectral Data
We 9 – 12 am	Review	Lecture 3, Lab 2
	Lecture 4	Sounding with broad band and hyperspectral IR
	Quiz 1	
We 2 – 5 pm	Lab 3	Investigations with Imagers and Sounders
Th 9 – 12 am	Review	Quiz 1, Lecture 4, Lab 3
	Homework 2 (du	e Saturday)
	Lecture 5	Microwave and time continuous (geostationary)
		measurements
Th 2 – 5 pm	Lab 4	Investigations of clouds and moisture with MODIS,
		AIRS and AMSU
	Lab	Assign and start Student Projects
Fr 9 – 12 am	Review	Homework 1, Lecture 5, Lab 4
	Lab	Student Projects
Fr 2 – 5 pm	Lab	Student Projects
Sa 9 – 12 am	Lab	Student Presentations using MODIS, IASI, AIRS, SEVIRI
	Review	Homework 2
	Lecture 6	Summary
	Quiz 2	

Remote Sensing Schools have been held in Bologna, Italy (Sep 01) Rome, Italy (Jun 02) Maratea, Italy (May 03) Bertinoro, Italy (Jul 04) Andanes, Norway (Feb 06) Cape Town, South Africa (Apr 06) Krakow, Poland (May 06) Ostuni, Italy (Jun 06) Benevento, Italy (Jun 07) Sao Paulo, Brazil (Nov 07) Monteponi, Sardinia (Sep 08) Istanbul, Turkey (Oct 08) Perth, Western Australia (Feb 09) Sasso di Castalda, Italy (Jul 09). New Dehli, India (Feb 11),













Objective of School

An in depth explanation of methods and techniques used to extract information from environmental satellite data, with emphasis on the latest measuring technologies. The course will consist of lectures, laboratory sessions, group lab projects, homework and tests. The results from each of the group projects will be presented to the class by the participating students. English is the official language of the School. All provided material will be in English.

Lectures and Labs

Lectures and laboratory exercises emphasize investigation of high spatial resolution visible and infrared data (from MODIS and SEVIRI), high spectral resolution infrared data (from AIRS and IASI), and microwave sounding data (AMSU). Text for the classroom and a visualization tool for the labs are provided free; "Applications with Meteorological Satellites" is used as a resource text from <u>ftp://ftp.ssec.wisc.edu/pub/menzel/</u> and HYDRA is used to interrogate and view multispectral data in the labs from <u>http://www.ssec.wisc.edu/hydra/</u>. Homework assignments and classroom tests are administered to verify that good progress is being was made in learning and mastering the materials presented. Classroom size is usually between twenty and thirty students.



Applications with Meteorological Satellites is used as a resource text

It is available for free at *ftp://ftp.ssec.wisc.edu/pub/menzel/*

CHAPTER 1 - EVOLUTION OF SATELLITE METEOROLOGY

CHAPTER 2 - NATURE OF RADIATION *

CHAPTER 3 - ABSORPTION, EMISSION, REFLECTION, AND SCATTERING *

CHAPTER 4 - THE RADIATION BUDGET

CHAPTER 5 - THE RADIATIVE TRANSFER EQUATION (RTE) *

CHAPTER 6 - DETECTING CLOUDS *

CHAPTER 7 - SURFACE TEMPERATURE *

CHAPTER 8 - TECHNIQUES FOR DETERMINING ATMOSPHERIC PARAMETERS *

CHAPTER 9 - TECHNIQUES FOR DETERMINING ATMOSPHERIC MOTIONS

CHAPTER 10 - AN APPLICATION OF GEOSTATIONARY SATELLITE SOUNDING DATA

CHAPTER 11 - SATELLITE ORBITS

CHAPTER 12 - RADIOMETER DESIGN CONSIDERATIONS *

CHAPTER 13 - ESTABLISHING CLIMATE RECORDS FROM MULTISPECTRAL MODIS MEASUREMENTS

CHAPTER 14 - THE NEXT GENERATION OF SATELLITE SYSTEMS

CHAPTER 15 – INVESTIGATING LAND, OCEAN, AND ATMOSPHERE WITH MULTISPECTRAL

MEASUREMENTS *

* indicates chapters covered

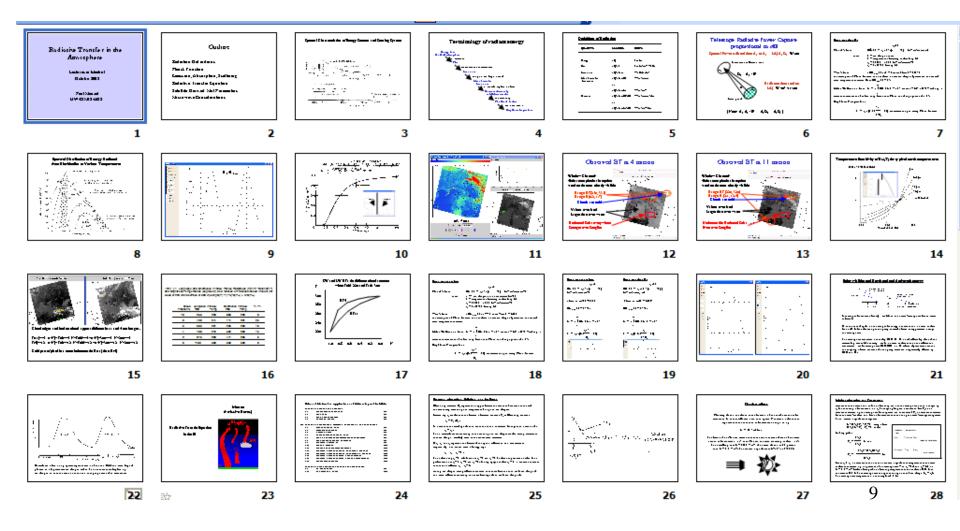
References, problems sets, and quizzes are included in the Appendices

Agenda includes material from Chapters 2, 3, 5, and 12

CHAPTER 2 - NATURE OF RADIATION

2.1	Remote Sensing of Radiation	2-1	
2.2	Basic Units	2-1	
2.3	Definitions of Radiation	2-2	
2.5	Related Derivations	2-5	
CHAPTER	3 - ABSORPTION, EMISSION, REFLECTION, AND SCATTERING		
3.1	Absorption and Emission	3-1	
3.2	Conservation of Energy	3-1	
3.3	Planetary Albedo	3-2	
3.4	Selective Absorption and Emission	3-2	
3.7	Summary of Interactions between Radiation and Matter	3-6	
3.8	Beer's Law and Schwarzchild's Equation	3-7	
3.9	Atmospheric Scattering	3-9	
3.10	The Solar Spectrum	3-11	
3.11	Composition of the Earth's Atmosphere	3-11	
3.12	Atmospheric Absorption and Emission of Solar Radiation	3-11	
3.13	Atmospheric Absorption and Emission of Thermal Radiation	3-12	
3.14	Atmospheric Absorption Bands in the IR Spectrum	3-13	
3.15	Atmospheric Absorption Bands in the Microwave Spectrum	3-14	
3.16	Remote Sensing Regions	3-14	
CHAPTER	5 - THE RADIATIVE TRANSFER EQUATION (RTE)		
5.1	Derivation of RTE	5-1	
5.10	Microwave Form of RTE	5-28	
CHAPTER 12 - RADIOMETER DESIGN CONSIDERATIONS			
12.3	Design Considerations	12-1	

Lectures are given with powerpoint presentations



Material includes equations

Planck's Law

 $c_{2}/\lambda T$ $B(\lambda,T) = c_{1}/\lambda^{5}/[e -1] \quad (mW/m^{2}/ster/cm)$ where $\lambda = \text{ wavelengths in cm}$ T = temperature of emitting surface (deg K) $c_{1} = 1.191044 \text{ x 10-5 (mW/m^{2}/ster/cm^{-4})}$ $c_{2} = 1.438769 \text{ (cm deg K)}$

Wien's Law $dB(\lambda_{max},T) / d\lambda = 0$ where $\lambda(max) = .2897/T$ indicates peak of Planck function curve shifts to shorter wavelengths (greater wavenumbers)with temperature increase. Note $B(\lambda_{max},T) \sim T^5$.

Stefan-Boltzmann Law $E = \pi \int_{0}^{\infty} B(\lambda,T) d\lambda = \sigma T^4$, where $\sigma = 5.67 \times 10^{-8} \text{ W/m}^2/\text{deg}^4$.

states that irradiance of a black body (area under Planck curve) is proportional to T^4 .

Brightness Temperature

$$T = c_2 / \left[\lambda \ln(\frac{c_1}{-+} + 1) \right]$$
 is determined by inverting Planck function
$$\frac{\lambda^5 B_{\lambda}}{\lambda^5 B_{\lambda}}$$

And some derivations,

$$\begin{split} I_{\lambda} \; = \; \epsilon_{\lambda}{}^{sfc} \; B_{\lambda}(T(p_s)) \; \tau_{\lambda}(p_s) + \Sigma \; \epsilon_{\lambda}(\Delta p) \; B_{\lambda}(T(p)) \; \tau_{\lambda}(p) \\ p \end{split}$$

The emissivity of an infinitesimal layer of the atmosphere at pressure p is equal to the absorptance (one minus the transmittance of the layer). Consequently,

$$\epsilon_{\lambda}(\Delta p) \ \tau_{\lambda}(p) \ = \ [1 \ \text{-} \ \tau_{\lambda}(\Delta p)] \ \tau_{\lambda}(p)$$

Since transmittance is an exponential function of depth of absorbing constituent,

$$\tau_{\lambda}(\Delta p) \tau_{\lambda}(p) = \exp \left[\begin{array}{cc} -\int & k_{\lambda} q \ g^{-1} \ dp \right] * \exp \left[\begin{array}{cc} -\int & p \\ \int & k_{\lambda} q \ g^{-1} \ dp \right] = \tau_{\lambda}(p + \Delta p)$$

$$p \qquad \qquad o$$

Therefore

$$\epsilon_\lambda(\Delta p) \; \tau_\lambda(p) \; = \; \tau_\lambda(p) \; - \; \tau_\lambda(p + \Delta p) \; = \; - \; \Delta \tau_\lambda(p) \; .$$

So we can write

$$\begin{split} I_\lambda \ = \ \epsilon_\lambda{}^{sfc} \ B_\lambda(T(p_s)) \ \tau_\lambda(p_s) - \Sigma \ B_\lambda(T(p)) \ \Delta \tau_\lambda(p) \ . \\ p \\ \end{split}$$
 which when written in integral form reads

$$I_{\lambda} = \epsilon_{\lambda}^{sfc} B_{\lambda}(T(p_s)) \tau_{\lambda}(p_s) - \int_{0}^{p_s} B_{\lambda}(T(p)) \left[d\tau_{\lambda}(p) / dp \right] dp.$$

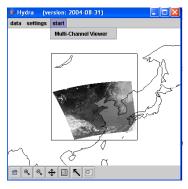


HYperspectral viewer for Development of Research Applications - HYDRA MODI

MSG, GOES

Freely available software For researchers and educators Computer platform independent Extendable to more sensors and applications Based in VisAD (Visualization for Algorithm Development) Uses Jython (Java implementation of Python) runs on most machines 512MB main memory & 32MB graphics card suggested on-going development effort

Rink et al, BAMS 2007



MODIS, AIRS, IASI, AMSU, CALIPSO

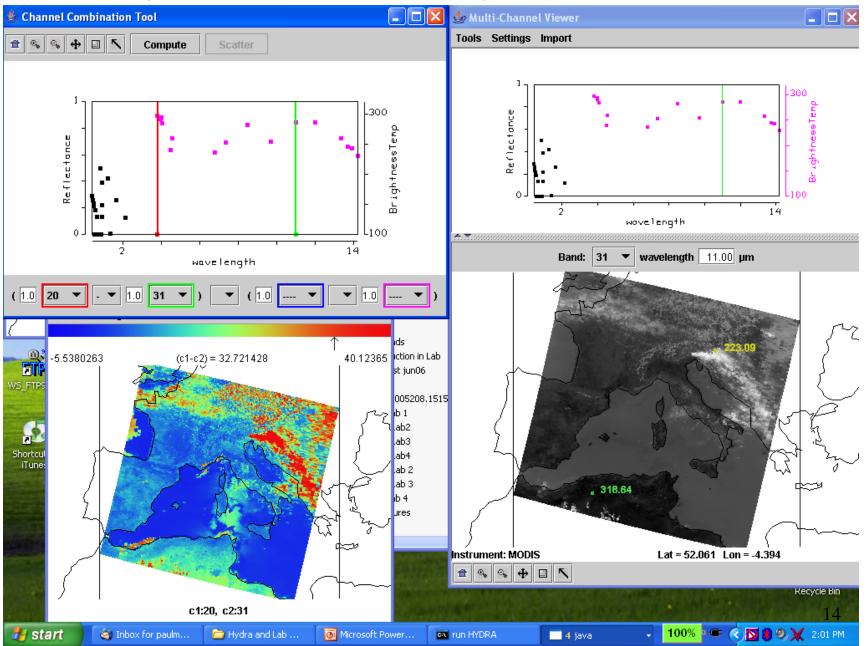
Developed at CIMSS by Tom Rink Tom Whittaker Kevin Baggett

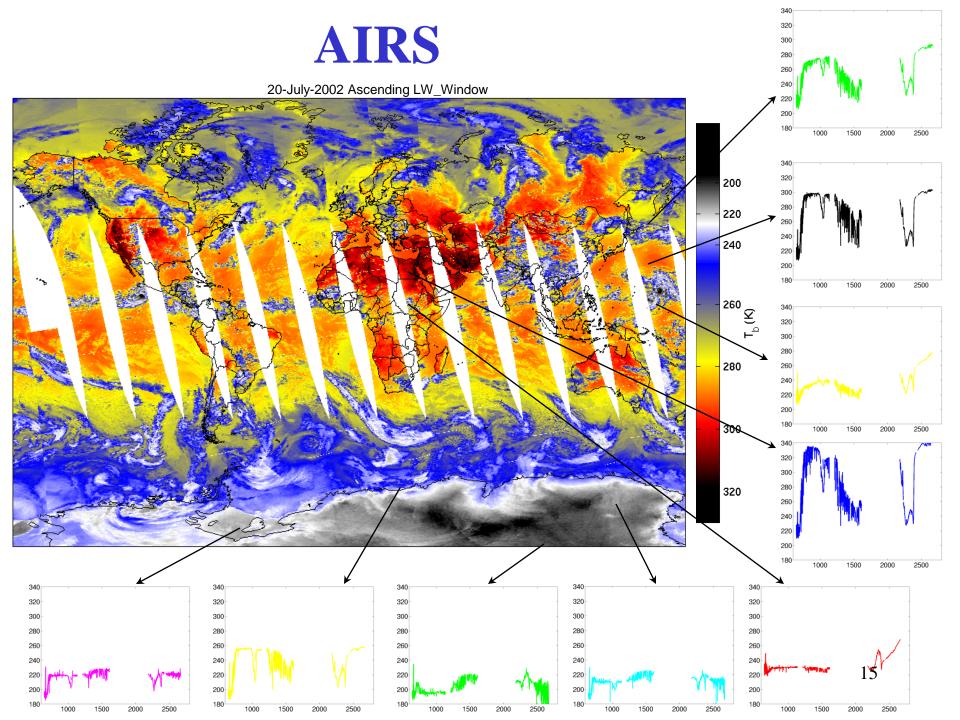
With guidance from Paolo Antonelli Liam Gumley Paul Menzel Allen Huang



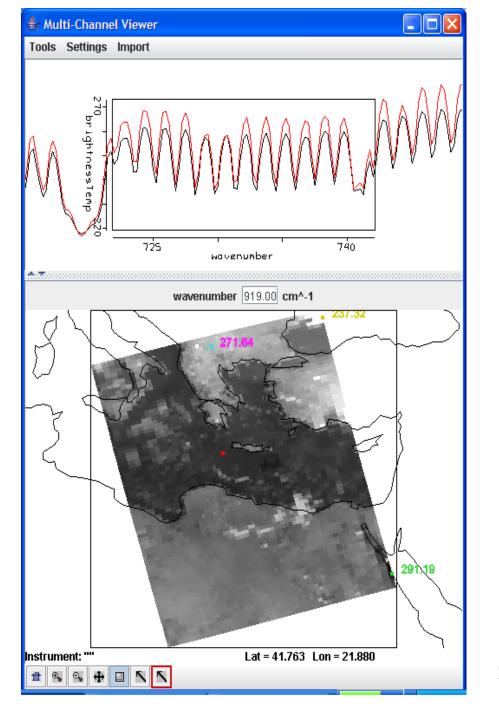
http://www.ssec.wisc.edu/hydra/

Viewing remote sensing data with HYDRA



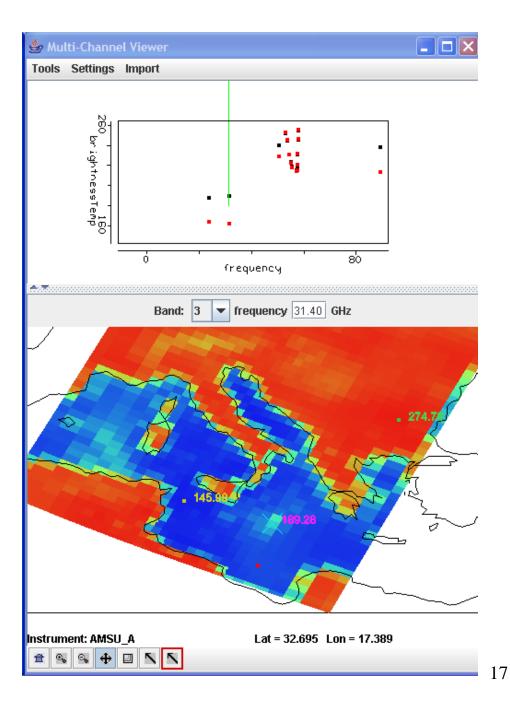


IASI



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Access to visualization tools and data

For hydra http://www.ssec.wisc.edu/hydra/

For MODIS data and quick browse images http://rapidfire.sci.gsfc.nasa.gov/realtime

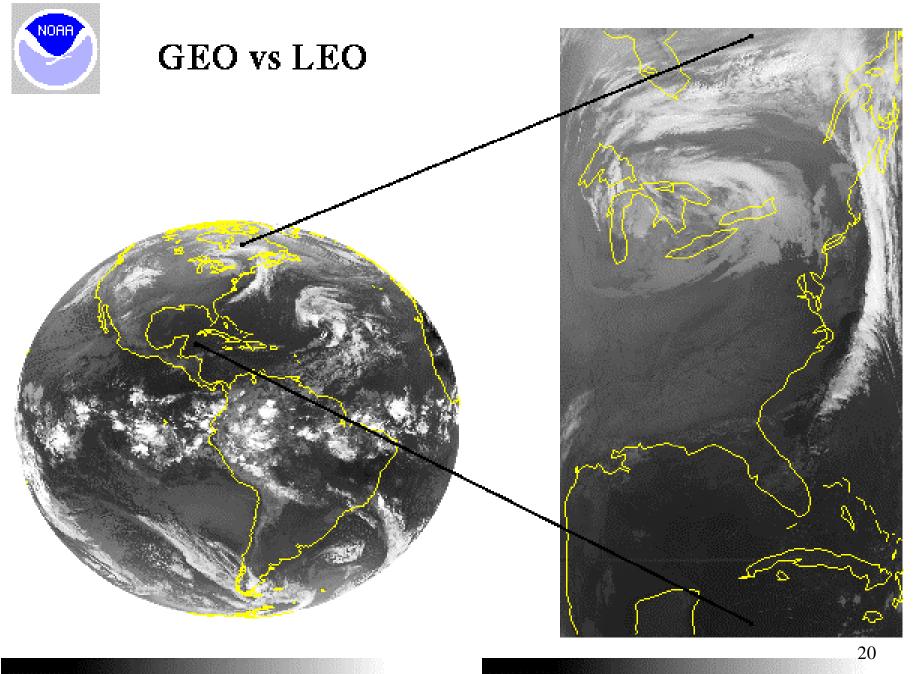
For MODIS data orders http://ladsweb.nascom.nasa.gov/

For AIRS data orders http://daac.gsfc.nasa.gov/

Orbits and Instruments

Lectures in Brienza 18 Sep 2011

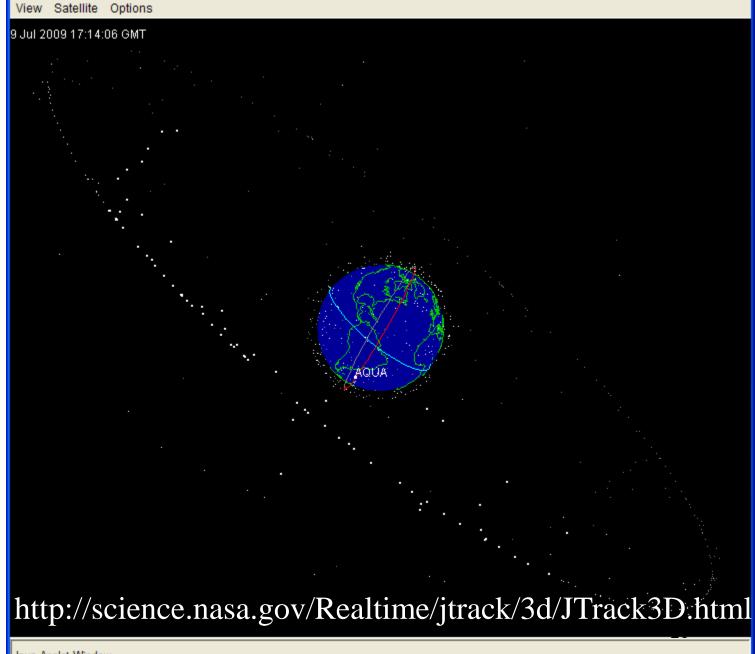
Paul Menzel UW/CIMSS/AOS



GOES-8 IMAGER 12UTC 02APR98

NOAA-12 AVHRR 12UTC 02APR98





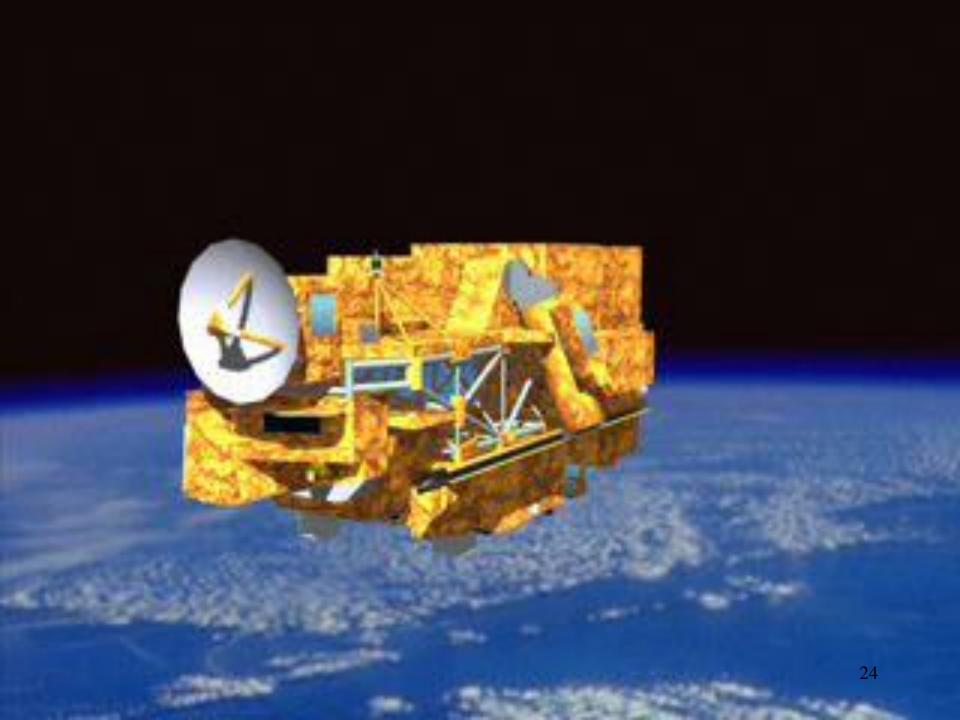
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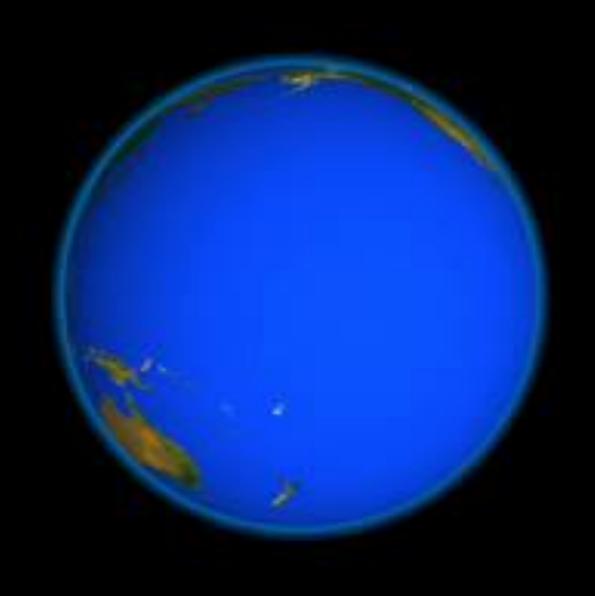
Java Applet Window

🍰 JTrack - 3D

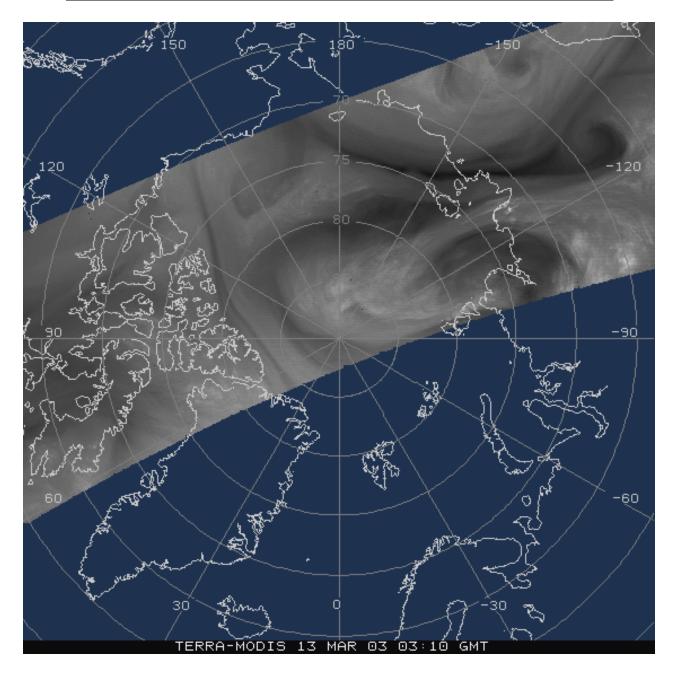






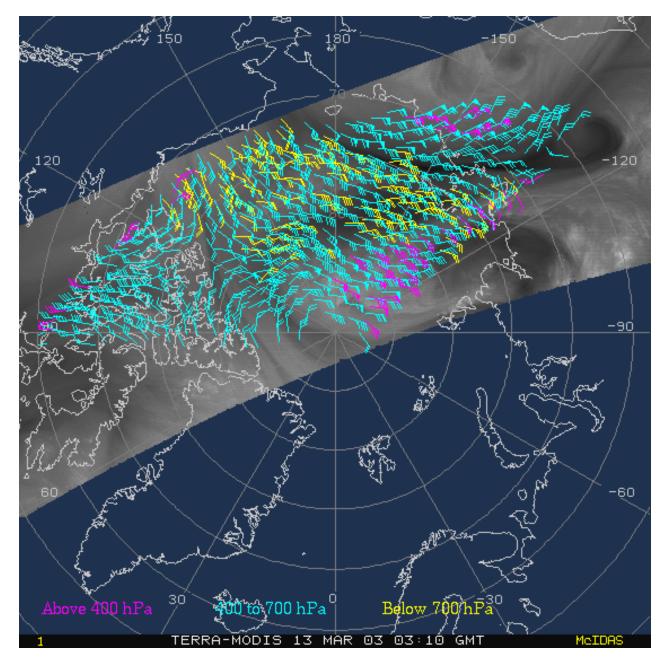


Leo coverage of poles every 100 minutes



26

Tracking Polar Atmospheric Motion from Leo Obs



27

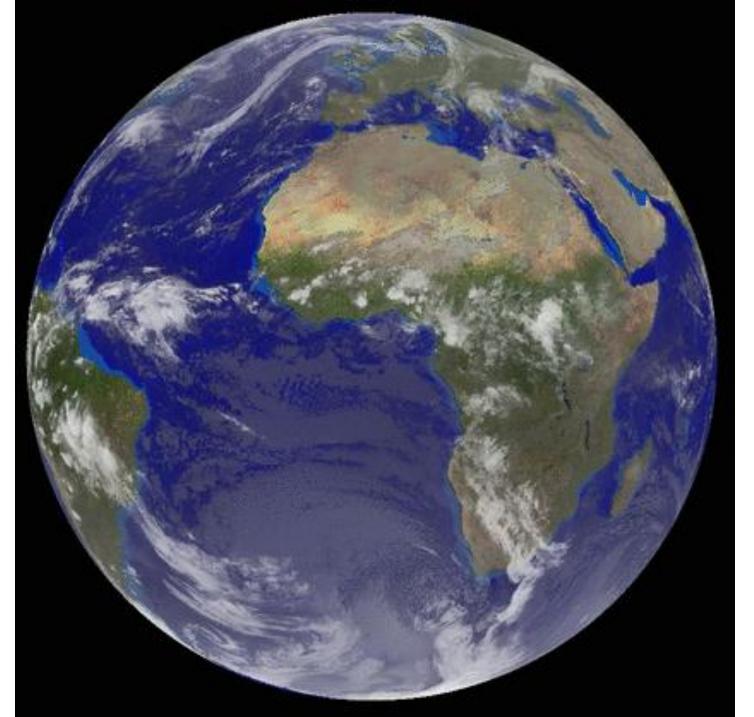
Getting to Geostationary Orbit



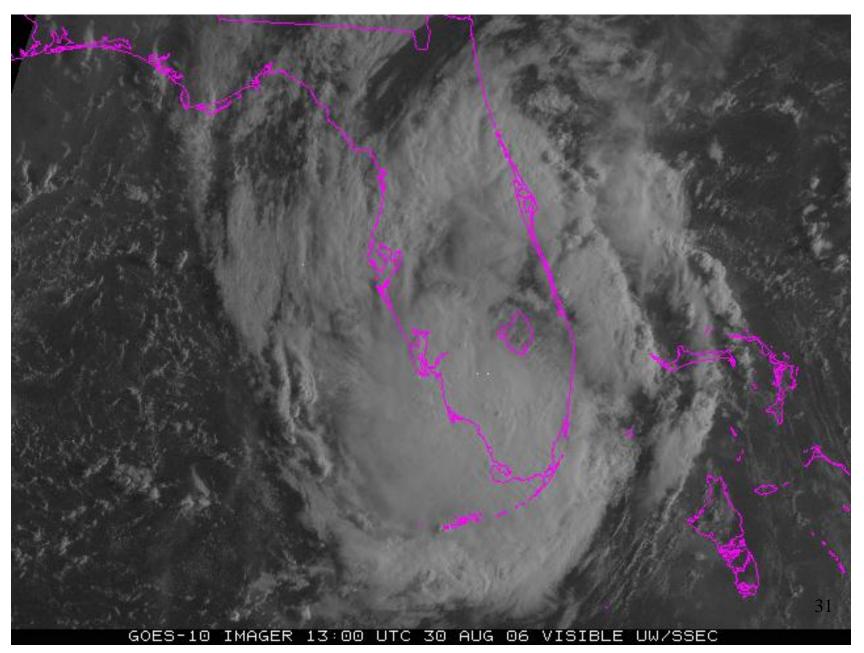
Observations from geostationary orbit



"the weather moves - not the satellite" Verner Suomi



One minute imaging over Florida



77 PBCJ



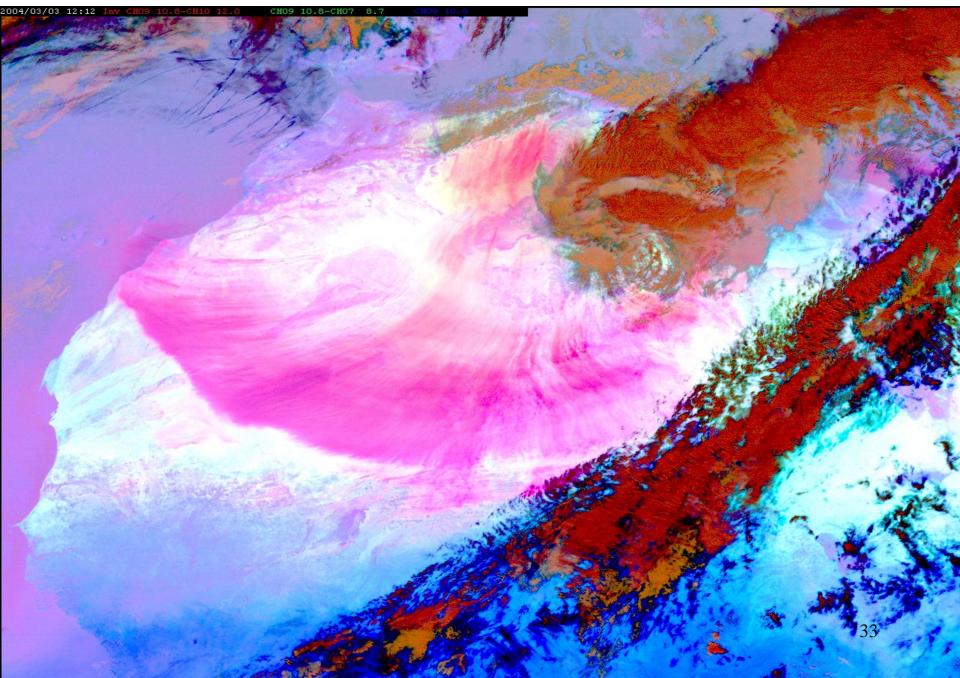


86⁶936 78 4200

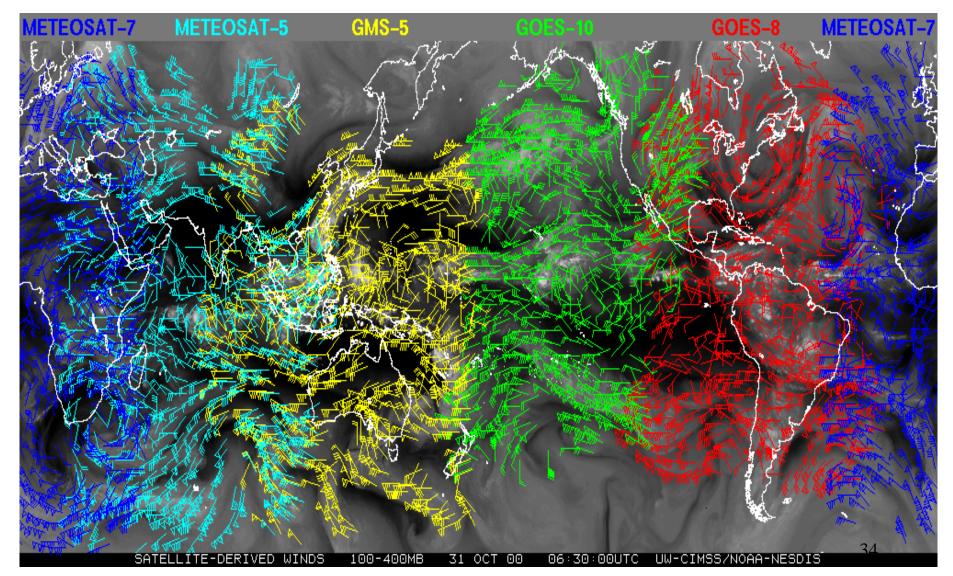
SURFACE PLOT - 12 UTC

GOES-12 IMAGER - VISIBLE (CH 01) - 11:45 UTC 28 AUG 2005 - CIMSS

SEVIRI sees dust storm over Africa



Five geos are providing global coverage for winds in tropics and mid-lats



<u>Comparison of geostationary (geo) and low earth orbiting (leo)</u> <u>satellite capabilities</u>

Geo

observes process itself (motion and targets of opportunity)

repeat coverage in minutes $(\Delta t \le 30 \text{ minutes})$

full earth disk only

best viewing of tropics

same viewing angle

differing solar illumination

visible, IR imager (1, 4 km resolution)

one visible band

IR only sounder (8 km resolution)

filter radiometer

diffraction more than leo

Leo

observes effects of process

repeat coverage twice daily $(\Delta t = 12 \text{ hours})$

global coverage

best viewing of poles

varying viewing angle

same solar illumination

visible, IR imager (1, 1 km resolution)

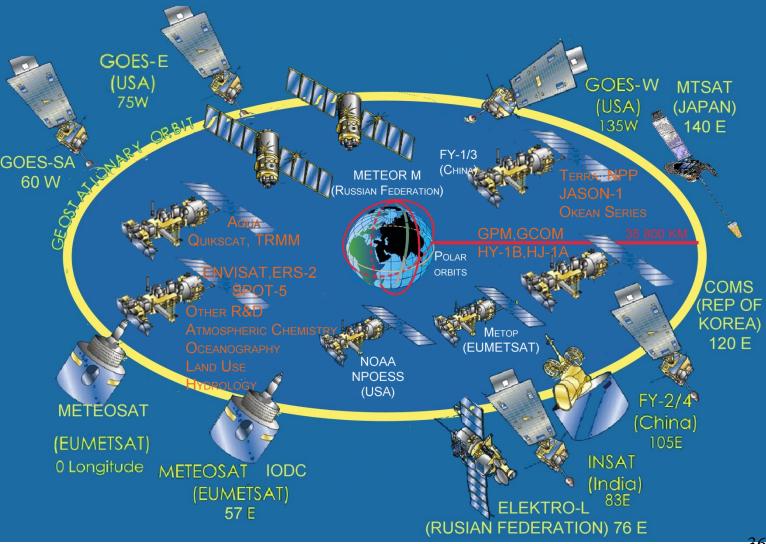
multispectral in visible (veggie index)

IR and microwave sounder (1, 17, 50 km resolution)

filter radiometer, interferometer, and grating spectrometer

diffraction less than geo

Space-Based component of the Global Observing System (GOS)



Leo Observations

Terra was launched in 1999 and the EOS Era began

MODIS, CERES, MOPITT, ASTER, and MISR reach polar orbit

> Aqua and ENVISAT followed in 2002

MODIS and MERIS to be followed by VIIRS AIRS and IASI to be followed by CrIS AMSU leading to ATMS



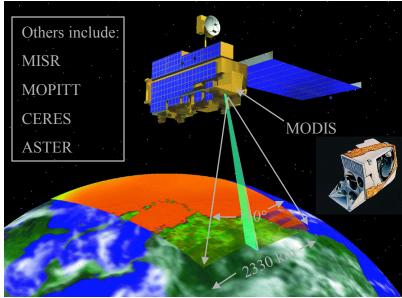


Launch of EOS-Terra (EOS-AM) Satellite - A New Era Begins





Launch date: December 18, 1999, 1:57 PT Earth viewdoor open date: February 24, 2001



MODIS instrument Specifications:

Bands 1-2 (0.66,0.86 μm): 250 m

Bands 3-7 (0.47, 0.55, 1.24, 1.64, 2.13 μm): 500 m

Bands 8-36: 1 km

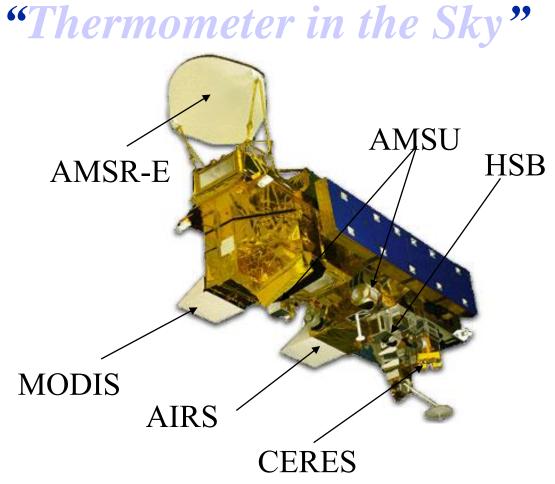
Allen Chu/NASA GSFC



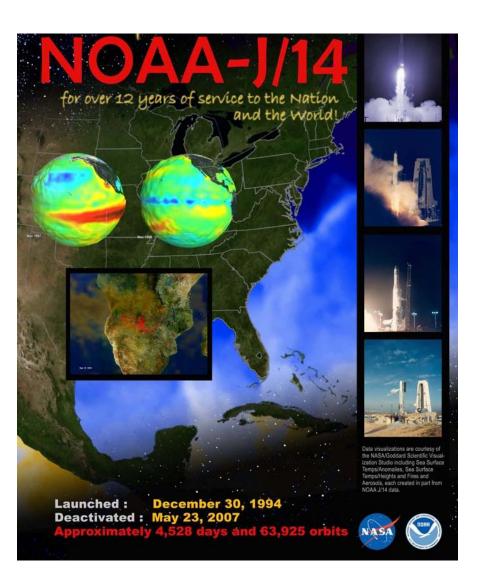
Followed by the launch of EOS-Aqua (EOS-PM) Satellite





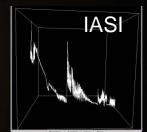


Launch date: May 4, 2002, 2:55 PDT Earth view door open date: June 25, 2002



Joint Polar System

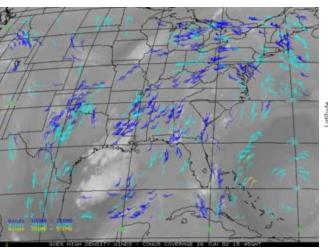
Welcome METOP Congratulations ESA / EUMETSAT



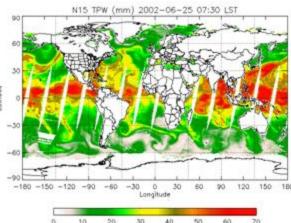
MetOp-A Launch on 19 October, 16h28 UTC Soyuz 2-1a, Baikonour

Atmospheric Products: Examples

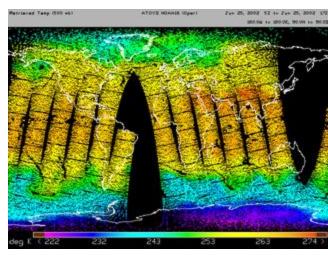
Winds



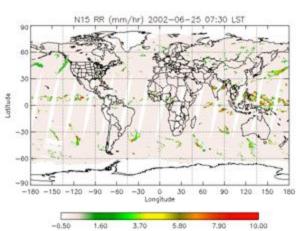
Total Water Vapor



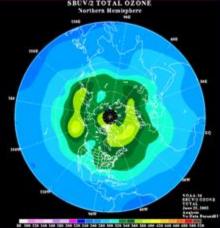
Temperature 500 mb



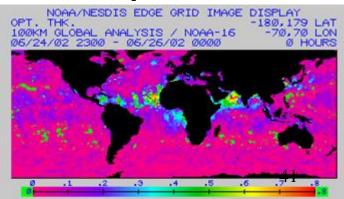
Rain Rate



Ozone

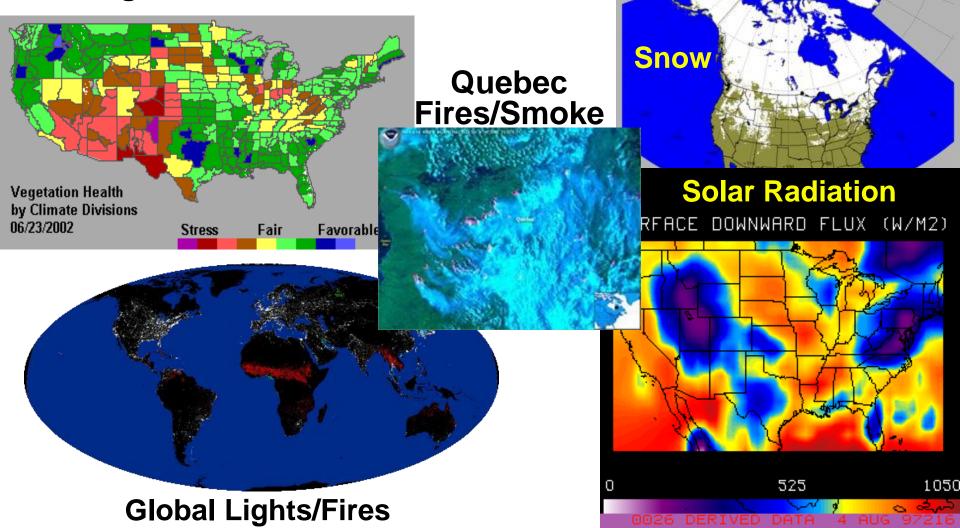


Aerosol Optical Thickness



Land Surface Products: Examples

Vegetation Health



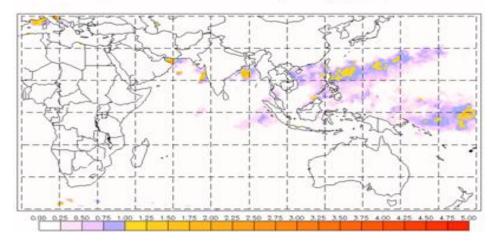
Ocean Products: Examples

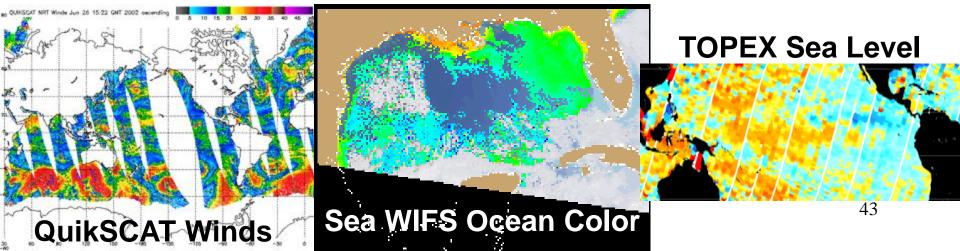
(enterregional Indicate sec-los)

SST Anomalies

Hot Spots: Potential Coral Bleaching

NOAA/NESDIS 50km SST - Maximum Monthly Climatology (C), 6/24/2002





Remote Sensing Advantages

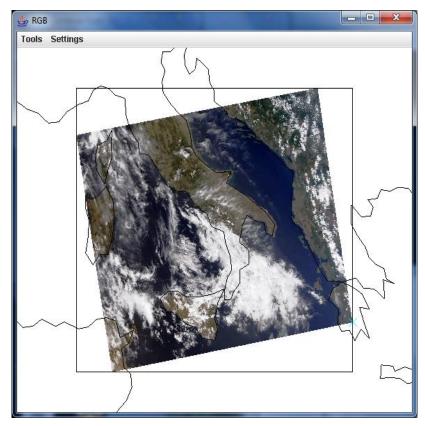
* provides a regional view

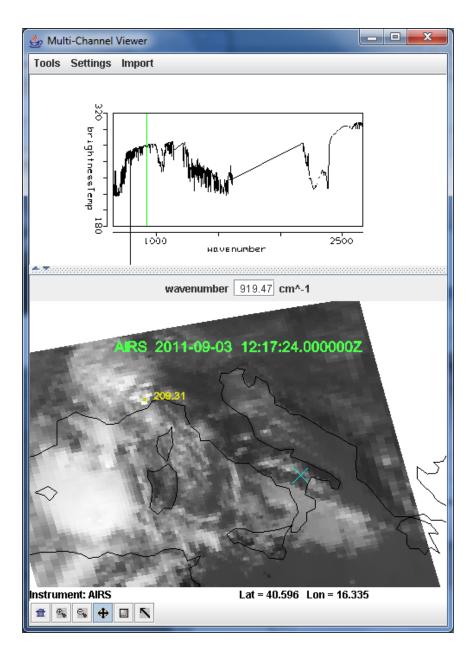
* enables one to observe & measure the causes & effects of climate & environmental changes (both natural & human-induced)

- * provides repetitive geo-referenced looks at the same area
- * covers a broader portion of the spectrum than the human eye
- * can focus in on a very specific bandwidth in an image
- * can also look at a number of bandwidths simultaneously
- * operates in all seasons, at night, and in bad weather

Welcome to a Short Course in Remote Sensing







Brienza Short Course in Remote Sensing 18 – 24 Sep 2011

