

Name: \_\_\_\_\_

1. Concerning Planck's radiation law:

(T) (F) The area under the Planck curve for isotropic radiation  $\iint B_\lambda d\lambda d\Omega$  has units of energy/ time/ area.

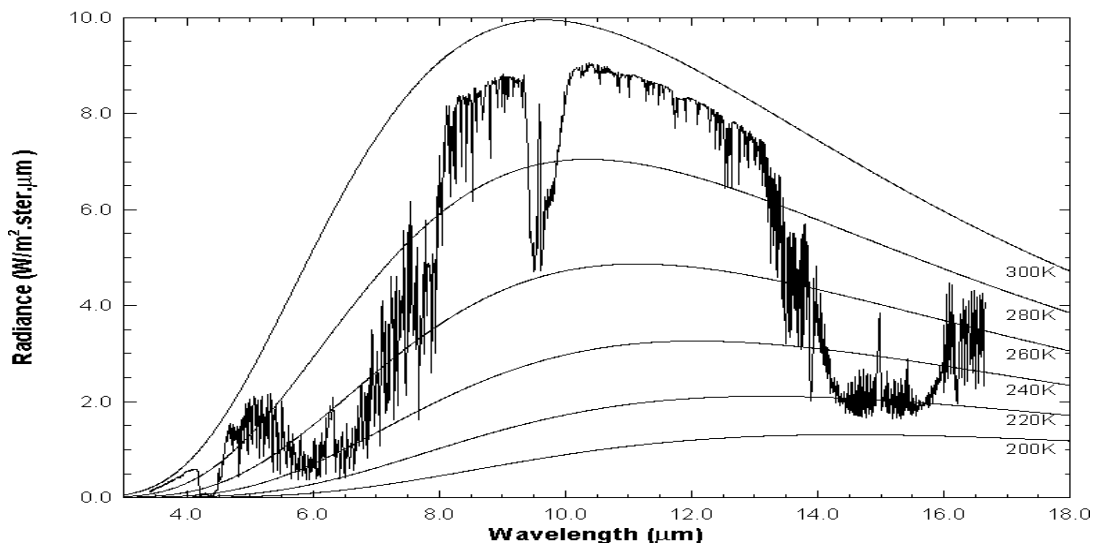
(T) (F) Wien's law can be derived by differentiating the Planck function  $B_\lambda$  with respect to temperature and equating the result with zero.

(T) (F) Radiance from a hotter source is greater at all wavelengths than the radiance from a cooler source (eg. if  $T_1 > T_2$  then  $B_\lambda(T_1) > B_\lambda(T_2)$  for all  $\lambda$ ).

(T) (F) The maximum of  $B_\lambda$  (radiance = irradiance per unit wavelength) versus  $\lambda$  (wavelength) is proportional to temperature to the fourth power,  $T^4$ .

(T) (F) For large wavelengths  $B_\lambda$ (radiance) is directly proportional to temperature.

2. Concerning the earth emitted radiances



(T) (F) The spectral region 10.5 to 12 μm is in a window region of the atmosphere.

(T) (F) The absorption band centers for H<sub>2</sub>O, O<sub>3</sub>, and CO<sub>2</sub> appear near 15, 9.7, and 6.5 μm respectively.

(T) (F) The tropopause temperature for this scene is roughly 220 K.

(T) (F) For a clear scene, the brightness temperature BT(13.3 μm) is typically colder than BT(13.9 μm).

(T) (F) 1% increase in temperature produces a larger percentage increase in radiance at 4 μm than 11 μm.